



Craters of the Moon National Monument and Preserve

Geology of Craters of the Moon



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Executive Summary

Craters of the Moon National Monument was established in 1924 to protect the geologic features of a small portion of southern Idaho's Great Rift volcanic rift zone. It was greatly expanded in November of 2000 and now covers approximately 750,000 acres (~300,000 hectares) or about 1,100 square miles (~2,800 square km). Craters of the Moon National Monument and Preserve now contains almost all of the Great Rift, the best-developed example of a volcanic rift zone on the Eastern Snake River Plain (ESRP). See figures 1 & 2 for setting and location. The Monument lies within the Snake River Basin-High Desert (Omernik, 1986) and is dominated by 3 geologically young (Late Pleistocene-Holocene) lava fields that lie along the Great Rift. The Great Rift varies in width between approximately 1 and 5 miles (1.6 and 8 km). It begins north of the Monument, about 6 miles (9.6 km) from the topographic edge of the Snake River Plain, in the vent area of the Lava Creek flows located in the southern Pioneer Mountains (Kuntz, et al, 1992). The Great Rift extends southeasterly from the Lava Creek vents for more than 50 miles (80 km) to at least as far as beneath Pillar Butte on the Wapi lava field (Kuntz, et al, 1982). The Great Rift volcanic rift zone is a belt of open cracks, eruptive fissures, shield volcanoes, and cinder cones.

The Craters of the Moon (COM) lava field is the northernmost and largest of the 3 young lava fields. Kings Bowl lava field is the smallest and lies between COM lava field and the Wapi lava field located on the southern end of the Great Rift. The other areas of the Monument, located either between these 3 young lava fields or surrounding them, are made up of Pleistocene age pahoehoe and a'a flows, near-vent tephra deposits, cinder cones, lava cones, and shield volcanoes (Kuntz, et al, 1988). These older areas are mantled with loess deposits (windblown silt) and in some places by eolian sand. Longitudinal sand dunes are prominent features surrounding the southern end of the Wapi

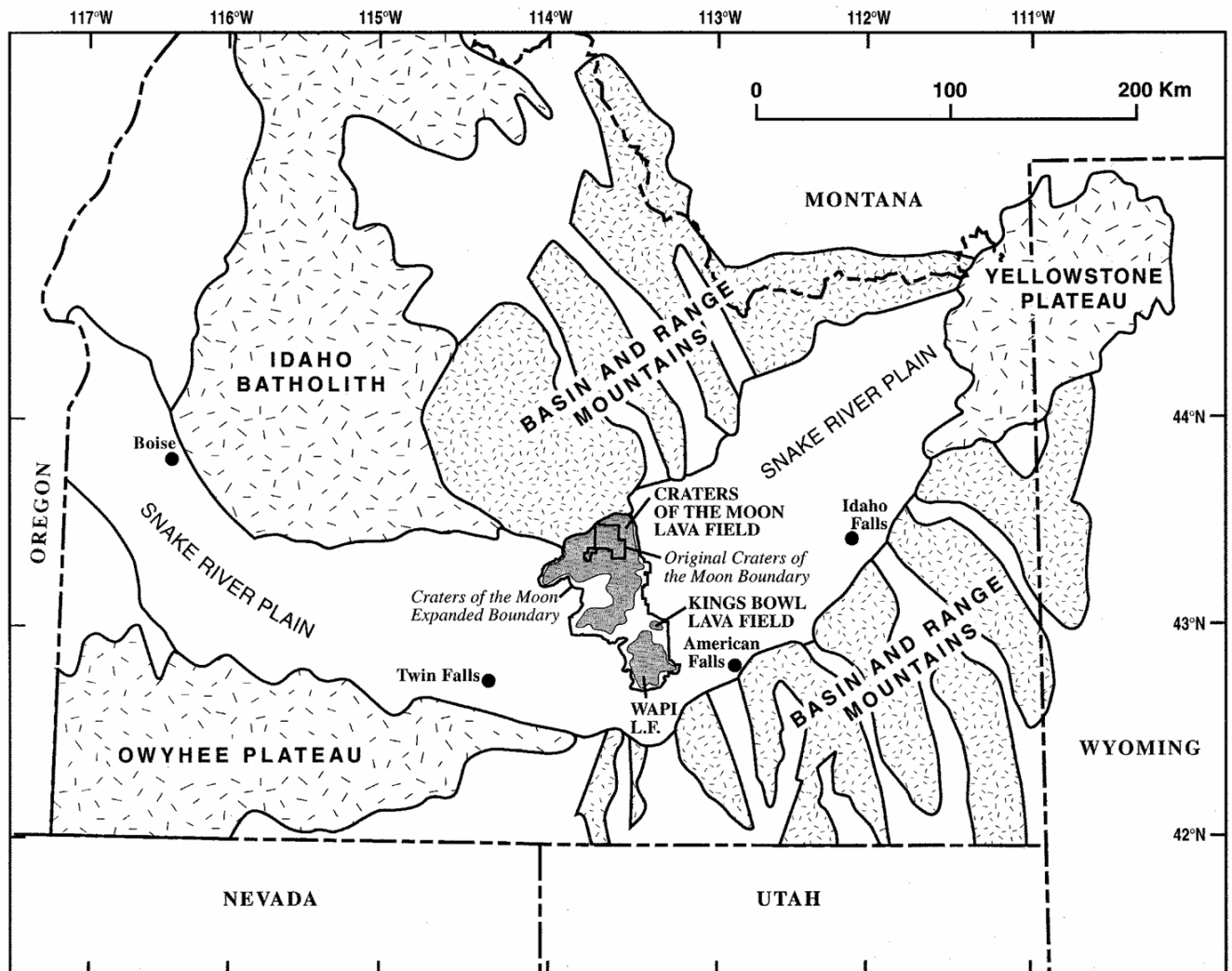
lava field (Greeley and King, 1977). During the Holocene (last 10,000 years), the highest volcanic activity of any of the eastern Snake River Plain (ESRP) basaltic rift systems was exhibited by these 3 lava fields associated with the Great Rift (Hughes, et al, 1999).

The COM lava field is the largest dominantly Holocene basaltic lava field in the lower 48 states (Kuntz, et al, 1992); it covers 618 mi² (1,600 km²). COM lava field is a composite field made up of at least 60 lava flows and 25 tephra cones. It has 8 eruptive fissure systems that are aligned along the northern part of the Great Rift (Kuntz, et al, 1992). The COM lava field has a tremendous diversity of volcanic features, with nearly every type of feature that is associated with basaltic systems (Hughes, et al, 1999). Unlike most of the ESRP, where the basalts are predominantly diktytaxitic olivine tholeiites (or more simply-- olivine basalts) associated with small monogenetic shield volcanoes (Hughes, et al, 1999), the basalt deposits in the COM lava field exhibit a wide range of chemical compositions. Though the COM lava flows are believed to have similar parent magma to the volcanoes in the rest of the Plain, their varied compositions are due to crustal contamination from assimilating older rocks or from crystal fractionation (Kuntz, et al, 1986). The COM lava field formed during at least 8 major eruptive periods over the past 15,000 years in contrast to most of the other lava fields on the ESRP that represent single eruptive events.

Kings Bowl lava field formed about 2,200 years ago during a single burst of eruptive activity that may have lasted as little as six hours (Kuntz, et al, 1992). Kings Bowl has a central eruptive fissure set that is about 4 miles long, which is flanked by 2 subparallel sets of non-eruptive fissures. The bowl that the field takes its name from is a phreatic explosion pit 280 feet (85 m) long, 100 feet (30 m) wide, and 100 feet (30 m) deep, caused by lava coming in contact with groundwater producing a steam explosion (Fig. 3). Adjacent to the bowl is an

outstanding example of a lava lake with well-developed levees. The crust of the lake

Figure 1. Regional setting and location of Craters of the Moon National Monument and Preserve



was broken by many of the blocks ejected by the phreatic explosion. The interior of this lake was still molten and oozed up through the holes punched in its crust, resulting in a large number of squeeze-up mounds of gas-charged lava (Hughes, et al, 1999). Many of the squeeze-ups look like mushrooms or candy kisses. There is also a plume of ash or tephra blanket on the east side of the pit that resulted from the explosion. Fissure caves, such as Crystal Ice Cave and Creons Cave lie along the Great Rift at Kings Bowl. At South Grotto, the rift may be passable to

a depth of 650 feet (198 m) below the surface (Earl, 2001). Feeding dikes drain

back features, and spatter cones can all be seen along the Great Rift at Kings Bowl.

Wapi and Kings Bowl lava fields are identical in age (approximately 2,200 years old), i.e., within the limits of analytical error (Hughes, et al, 1999). The Wapi lava field is a classic shield volcano with a flattened dome shape. Kuntz, et al, (1992) believe that the Wapi lava field began as a fissure eruption, but with prolonged activity developed a

sustained eruption from a central vent complex, which produced the low shield volcano seen today. The vent complex is made up of 5 major and 6 smaller vents, with the vents being steep-sided circular depressions typically about 300 feet (90 m) in diameter and 30 feet (9 m) deep (Kuntz, et al, 1992). Rising about 60 feet (18 m) above the south side of the largest vent is a mass of agglutinate and layered flows known as Pillar Butte. Medial and distal parts of the lava field are mostly composed of tube-fed pahoehoe flows. Pressure plateaus, flow ridges, and collapse depressions characterize the margins of the field where local relief can be over 30 feet or 9 meters (Kuntz, et al, 1992). Greeley, 1971, reported that the only known dribblelet spires in the continental U.S. occur on the flows associated with Pillar Butte. Now however, dribblelet spires are known to occur at least also in Diamond Craters in Oregon. The spires found in the Wapi lava field average 12 feet (3.6 m) high and 5 feet (1.5 m) in diameter. They consist of imbricated rounded slabs of lava that range from 5 inches by 5 inches (12.7 cm by 12.7 cm) by 1 inch (2.5 cm) thick to 12 inches by 9 inches (30.5 cm by 22.9 cm) by 4 inches (10.2 cm) thick (Greeley, 1971). These bizarre spires are a type of hornito.

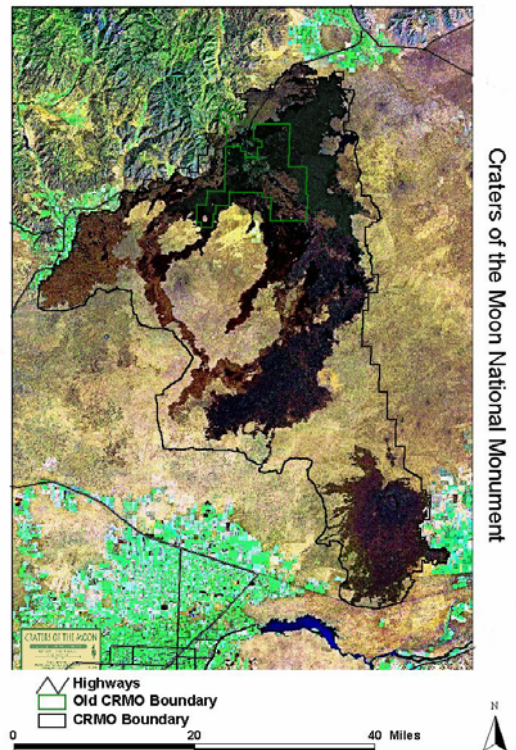


Figure 2. Craters of the Moon (CRMO)



Figure 3. View of King's Bowl and Great Rift.

History of Geologic Exploration

Native Americans have visited the area of Craters of the Moon National Monument for thousands of years. They were potential witnesses to at least the last 3 eruptive periods of the COM lava field and for the formation of both the Wapi and Kings Bowl lava fields. Scientific studies began in 1901, when Israel Russell, United States Geological Survey (USGS), came to investigate south-central Idaho. Russell wrote the first

scientific account of the region called "Cinder Buttes" (Russell, 1902). His intense interest in the volcanic bombs that he discovered brought him back again the following year (Russell, 1903). Starting in 1910, S.A. Paisley, who later became the first Custodian of the Monument made numerous trips into the area and Era Martin, a local resident, discovered and marked many of the caves and water holes with stone monuments (Stearns, 1928).

In 1921, Robert Limbert, a taxidermist from Boise, Idaho, visited the area and published an account of his trip in *National Geographic* (Limbert, 1924). In the same year, O. Meinzer, Chief of USGS Division of Ground Water, and Harold Stearns, USGS, visited the area. In 1923 Stearns, accompanied by F.E. Wright of the Carnegie Institution made a trip to the area and published a description in *Geographical Review* (Stearns, 1924). The National Park Service (NPS) requested Stearns to submit a report describing the area, delineating boundaries, and stating the reasons that would justify its preservation as a National Monument and on May 2, 1924 President Calvin Coolidge proclaimed the original Craters of the Moon National Monument. In 1925, the first topographic map of the Monument was made by M.J. Gleissner, USGS (Stearns, 1928). In the fall of 1926, Stearns again returned to the Monument and spent a month mapping the geology and describing the features. He submitted a report to the USGS, but only portions of it were published (Stearns, 1928; Stearns, et al, 1938). The 1938 publication contained the first generalized geologic map of the monument.

Since these early geologic investigations, numerous additional studies have been conducted to workout the regional geology and structural setting, the source of the volcanism, the petrology and mineralogy of the lavas and underlying rocks, and the chronology of the lava flows. The 1950's and 60's produced many reports speculating on the structure of the western Snake River Plain (Malde, 1959; Malde and

Powers, 1962; Malde, et al, 1963; Hill, 1963). In the 1970's detailed petrologic and geochronologic studies were initiated and research began on the geochemistry of the basalts and underlying rhyolites of the plain. Simultaneously, dating began using both radiocarbon and K-Ar (potassium-argon) techniques to determine the absolute age of the lava flows. Mineralogical studies at this time focused on the Blue Dragon flow in the Monument and on the unusual mineral deposits found in some of the lava tubes and pits (e.g.: Armstrong, et al, 1975; Bullard, 1970; Faye and Miller, 1973; Malde and Cox, 1971, Peck, 1974). Christiansen and McKee (1978) published the first report of a mantle plume theory for the formation of the Snake River Plain. From the 1980's to the present geologists concentrated on refining the petrologic characterization and absolute dating of the rocks and on magmatic and eruptive models to explain the regional volcanism. Kuntz, et al, (1982) dated lava flows within the monument and (Kuntz, et al, 1986) provided significant information regarding the source, volume and periodicity of the basaltic eruptions. Smith and Braile (1993) described the space-time evolution of the Yellowstone-Snake River Plain volcanic system. Geophysical techniques, such as tomography (imaging based on P-Wave velocity structure), were used to determine the subsurface geology of the plain (Humphreys, et al, 2000).

Geologic Setting

The Monument lies on the eastern portion of the Snake River Plain (SRP), which is a topographic low in southern Idaho that is 30 to 60 miles (50 to 100 km) wide and extends 360 miles (600 km) from Oregon to the Yellowstone Plateau. The plain is bounded by highlands to the north and south (Fig. 1). The Eastern Snake River Plain (ESRP) consists of broad flat basalt flows and thin discontinuous sedimentary deposits that together have a total thickness of ~0.6 to 1.2 miles or ~1 to 2 km (Doherty, et al, 1979). Magnetic polarity determinations and recent radiometric studies (Champion, et al, 1988; Kuntz, et al, 1992) indicated that most of the surface flows were erupted during the

Brunhes Normal-Polarity Chron, and thus are younger than 780 ka (ka = thousand years ago). Data from wells that penetrate the ESRP to depths as great as 2 miles (3,500 m) show that the lava flow and sediment sequence is 0.6 to 1.2 miles (1 to 2 km) thick throughout most of the plain (Embree, et al, 1982). Drilling and field studies (Doherty, et al, 1979; Embree, et al, 1982; Morgan, et al, 1984) show that the basalt-sediment sequence is underlain by rhyolitic lava flows, ignimbrites (rock formed by the widespread deposition and consolidation of ash flows), and pyroclastic deposits (formed by volcanic explosion or aerial expulsion from a volcanic vent).

The structure of the SRP varies greatly from west to east. The northwestern plain is believed to be a graben, or fault-bounded depression (Malde, 1959). The ESRP, which includes the Monument, is less clearly defined. The earliest studies by Kirkham (1931) hypothesized that the plain was a downwarp, which had been filled with basalt both during and following the downwarping process. He believed that extrusion of lava was the major cause of subsidence and he supported this hypothesis with evidence of volcanic layers that gently dip toward the center of the plain.

Some have suggested that the volcanism and structure of the plain are the result of an eastward propagating rift (Myers and Hamilton, 1964; Hamilton, 1987), and transform fault boundaries across basin and range faults (Christiansen and McKee, 1978). Or the plain may simply be related to a preexisting crustal weakness, i.e., the structure of the Precambrian basement in southern Idaho (Eaton, et al, 1975). A more catastrophic explanation for the formation of the plain is related to a hypothesized meteorite impact in southwestern Idaho. The impact is conjectured to have caused deep fractures in the earth's crust that initiated the eruption of flood basalts, which were followed by lower-volume outpouring of lava from the fractures (Alt and Hyndman, 1988).

Some of the most popular recent explanations for the formation of the ESRP have involved theories incorporating a deep mantle plume. Thick rhyolitic rocks encountered during drilling in the ESRP suggested that some portions of the plain represent filled rhyolitic-calderas, similar to the Henrys Fork (Island Park) Caldera in Idaho (Doherty, et al, 1979; Embree, et al, 1982; Morgan, et al, 1984). This information, along with some geophysical data, led many geologists to believe that the ESRP is the site of a northeasterly propagating system of rhyolitic volcanic centers. It is thought that the southwesterly movement of the North American Continent, caused by plate tectonics, has passed southern Idaho over a stationary mantle plume or hotspot. In turn, the mantle plume has caused rhyolitic and associated basaltic volcanism to develop across southern Idaho.

This mantle plume is believed to be the same heat source for the volcanic and hydrothermal activity in Yellowstone National Park (e.g.: Armstrong, et al, 1975; Brott, et al, 1981; Maley, 1987; Pierce and Morgan, 1992). Some deep mantle-plume advocates believe that the plume ascended from the core-mantle boundary, was progressively overridden by the North American Plate starting about 60 million years ago, and may be responsible for the Carlin gold deposits found in Nevada (Oppliger, et al, 1997). Along the same line, others hypothesize that the mineralization in the Carmack Group, found in the Yukon, is related to the Yellowstone hotspot of some 70 million years ago (Johnston, et al, 1996).

Regardless of when the hotspot originated, the ESRP records a progressively younger trend of rhyolitic eruptions to the northeast. Henry Heasler with the Yellowstone Volcano Observatory reports that there have been 142 massive blasts, catastrophic eruptions of huge volumes of rhyolitic magma, in the last 17 million years along the ESRP (Sparrow, 2003). These eruptions typically produced calderas 10-40 miles wide. Many of the

calderas overlapped and may be broken down into 7-13 volcanic centers. Although some of the mountain ranges that existed on the ESRP before the hotspot may have been blown away by the eruptions, it is more likely that they were swallowed up as the floor of the caldera sank during the violent explosions, thus producing the trough we see today (Smith and Siegel, 2000). Kuntz, et al, (1992) believe that the source of the material for the ESRP eruptions is lithospheric mantle with the melting being driven by plume upwelling and decompression-melting. In contrast to a deep mantle plume, more recent teleseismic (utilizing distant seismic events) studies led Smith and Siegel (2000) to believe that the root of the hotspot is only at a depth of about 125 miles (200 km). Humphreys, et al, (2000) envision convective rolls within the athenosphere or local upper-mantle convection instead of a deep mantle plume.

This recent seismic data also suggests that the Yellowstone hotspot left behind a slab of basalt 6-10 miles thick in the mid-crust and that it contains partial melt. Smith and Siegel (2000) figuratively describe this slab as representing the slag left in the bottom of the numerous magma chambers spawned by the hotspot. The region surrounding the ESRP continues to experience basin-and-range type faulting, which is stretching or pulling apart the crust. This crustal extension continues to uplift the mountain ranges, such as the Lost River Range where a magnitude 7.3 earthquake occurred in 1983. On the ESRP in the wake of the Yellowstone hotspot, where all of these hot rocks have been left behind, instead of producing mountain ranges, the tensional forces help to create decompression melting, which results in dike emplacement and periodic eruption of molten rock onto the surface. As long as these forces continue to act, more eruptions will eventually occur. It is estimated that the ESRP is made of 8,000 shield volcanoes and the typical volume erupted is 1.2 mi³ or 5 km³ (Kuntz, et al, 1992). Coalesced shield volcanoes and lava cones constitute >95%

of the total volume of basalt in the ESRP and the lava flows are dominantly of the tube-fed type (Kuntz, et al, 1992).

Several prominent rhyolitic domes lie along the ESRP in the Arco Desert and are visible from ten's of miles away. The tallest, Big Southern Butte, is a landmark and navigation aid visible from much of the Monument. It towers 2,500-ft (760m) over the ESRP and extends another 2,950-ft (900 m) into the subsurface. Big Southern Butte, one of the largest rhyolite domes in the world, has been dated at ~300 ka. Other buttes of the ESRP include East Butte (~600 ka), Middle Butte (~600 ka) and Cedar Butte (~400 ka). Current research indicates that these rhyolite domes formed through extreme fractionation of basaltic magma (McCurry, et al, 1999).

Basaltic Volcanism

The basaltic volcanism on the ESRP is localized in lava fields along volcanic rift zones. These zones are narrow belts, typically 3-12 miles (5-20 km) wide, composed of faults, grabens, eruptive and non-eruptive fissures, spatter cones, cinder cones, and low shield volcanoes (Kuntz, 1977a, 1977b; Kuntz, et al, 1992). Most volcanic rift zones are perpendicular to the long axis of the plain, and may be extensions of faults that bound basin-and-range mountains north and south of the plain (Kuntz, 1977b). Eight separate young basaltic lava fields (Craters of the Moon, Kings Bowl, Wapi, Shoshone, North Robber's, South Robber's, Hell's Half Acre, and Cerro Grande) can be identified in Landsat images of the ESRP (Lefebvre, 1977; Champion and Greeley, 1977; King, 1977). The largest of these is the Craters of the Moon (COM) lava field, which is also the largest basaltic lava field of dominantly Holocene age (less than 10,000 years) in the conterminous United States.

The COM lava field consists of more than 60 lava flows that cover an area of 618 mi² (1600 km²) and a volume of 7.2 mi³ (30 km³). The volcanic vents that supplied the lava flows of the COM lava field are aligned

along the northern part of the Great Rift. The Great Rift is approximately 53 miles (85 km) long and extends from the southern Pioneer Mountains southeastward through the Monument to Pillar Butte, located about 18 miles (30 km) northwest of American Falls. The Great Rift is the best example of a volcanic rift zone on the ESRP and can be divided into four sections. The northernmost section lies beneath the COM lava field. Just south of the COM lava field, the rift changes to an inactive open fissure system. The southern portions of the rift gave rise to the Kings Bowl lava field and to the Wapi lava field. Figure 4 is a view down the axis of the Great Rift in Kings Bowl.

The variety of lava flows and volcanic vents in the COM lava field represent a nearly complete range of the types of volcanic features formed by basaltic eruptions. Lava flows and volcanic landforms, such as cinder cones and spatter cones, are highly concentrated in the northern end of the COM lava field. Based on observations of active basaltic eruptions, geologists believe that these volcanic landforms are the result of combinations of distinct eruptive phases (Kuntz, et al, 1982).



Figure 4. Great Rift at South End of Kings Bowl (Note- people at top for scale).

Ejection of lava by both curtains of fire and fire fountains can produce spatter ramparts, cinder cones, mounds of cinder, and generate several kinds of volcanic bombs. The fire fountains that produced many of the Monument's cinder cones were probably >1,000 feet (300 m) high. Big Cinder Butte, the tallest cinder cone in the Monument is over 700 feet (210 m) high and may have had a fire fountain >1,500 feet (450 m) high.

Four kinds of bombs are found in the Monument, all of which started off as globs of molten rock thrown or ejected into the air. If the glob got twisted during its flight it is called a spindle bomb (Fig. 5) and typically range from a few inches to several feet in length. If the bomb was very tiny and

twisted it is called a ribbon bomb. When a glob of molten rock forms a crust as it flies through the air and the gases inside continue to expand and crack that crust, it is called a breadcrust bomb, because of its similarity to bread rising in an oven (Fig. 6). If Many eruptions in the COM lava field are believed to have begun with a phase characterized by a "curtain of fire". Curtains of fire are long lines of gas-charged lava erupting in low fountains. These curtains can extend for up to several miles, are generally 100-200 feet (30-60 m) tall, and can be sustained for hours and up to several days. As a curtain of fire continues to erupt, segments of the fissure begin to clog. This results in the same amount of lava being forced to erupt from a limited number of vents producing fire fountains.



Figure 5. Spindle Bomb



Figure 6. Breadcrust Bomb

The first landforms to develop from the lava fountains are cinder cones, which are accumulations of cinders (light fragments riddled with gas holes), volcanic bombs, spatter and lava flows that collect around

the bomb did not completely solidify during flight and flattened on landing, it is called a cow-pie bomb or is sometimes also referred to as a pancake bomb. Some cow-pie bombs in the Monument are over 10 feet long.

the vent forming a cone with a central crater. There are more than 25 major cinder cones within COM lava field. Prevailing winds from the west and southwest have caused a preponderance of downwind accumulation of cinders from many of the vents. This has resulted in an elongation of many cinder cones to the east or northeast, making them asymmetrical. The Great Rift volcanic rift zone is about 1.5 miles (2.5 km) wide in the COM lava field where many of the cinder cones are found. The cones are generally located along the outer margins of this zone. Cones on the western margin include (north to south) Grassy Cone, Silent Cone, Big Cinder Butte (Fig. 7), Echo Crater, the Sentinel and Fissure Butte. The eastern margin includes Sunset Cone, Paisley Cone, Half Cone, Broken Top, and the Watchman cinder cones.

After several hours, days, or weeks, a decrease in magma pressure and the amount of dissolved gases in the magma can produce a corresponding change in the output of lava. At this time, the amount of lava spraying from the vent(s) decreases and begins pouring out as lava flows. This phase may last many years and start and stop many times. The lava flows can vary greatly in temperature and composition from one vent to another and with time;



Figure 7. Big Cinder Butte

therefore, they also vary in viscosity. This creates fields of pahoehoe, slabby pahoehoe, a'a, and block lava flows.

The source of these lava flows can be from the same vent that formed a cinder cone, spatter rampart or spatter cone. When a cinder cone is the source of a flow, the lava burrowing through the side of the cone may breach the cone. This commonly results in a notch in the cone above the feeder fissure. The lava flow that breaches the cone may occur concurrently with the cinder cone development or it may occur long after the formation of the cone because of reactivation of the fissure system underlying it. In the COM lava field, reactivation of this sort probably occurred in vents of North Crater, Broken Top, and the Watchman cinder cones where the lava flows seem to be significantly younger than the cones. It is also common for lava flows to originate from portions of active rifts that have not previously undergone fountain-type eruptions. In this case, lava can flow directly out of the unobstructed vent and onto the landscape.

When the eruption of lava continues for a long time from an unobstructed vent, a large shield volcano can be produced. Shield volcanoes are gently sloping and have a flattened dome shape. Wapi lava field in the southern part of the Monument is an outstanding example of a shield volcano. Kuntz, et al, (1992) estimated from the calculated volume of basalt on ESRP and typical eruption volumes that the ESRP is made up of about 8,000 shield volcanoes. The largest shields are typically located near the center of the plain, suggesting that they overlie the central region of magma generation. Lava can travel great distances through lava tubes with very little loss of heat. Lava tubes and tube systems, therefore, facilitate the transport of lava over great distances. Some flows extend up to 30 miles or 48 km (Hughes, et al, 1999).

Lava is described by its physical appearance, which is largely determined by its composition, temperature, fluid and

crystal content, and the influence the surface and slope it is flowing down exert on it. Block lava (Fig. 8) has a surface of angular blocks and forms from very dense lava. The typical composition of block lava in the Monument is trachyandesite (an extrusive rock, intermediate in composition between trachyte and andesite). A'a has a rough, jagged, or clinkery surface. Pahoehoe has a smooth, ropy, or billowy surface (Fig. 9). Pahoehoe can be further broken down into several types. Shelly pahoehoe forms from highly gas-charged lava, often near vents or tube skylights, and contains small open tubes, blisters and thin crusts. Spiny pahoehoe forms from very thick and pasty lava and contains elongated gas bubbles on the surface that form spines. Spiny pahoehoe is the dominant form found in the Monument. Slabby pahoehoe is made up of jumbled up plates or slabs of broken pahoehoe crust. Both slabby and spiny pahoehoe are transition phases to a'a.

Lava tubes, which are hollow spaces beneath the surface of solidified lava flows, are formed by the withdrawal of molten lava after the formation of the surface crusts. Indian Tunnel in the northern part of the Monument has a 40-foot (12 m) high ceiling and is 800 feet (240 m) long (Fig. 10). Bear Trap Cave, which lies between COM and Kings Bowl lava fields is >10 miles (16 km) long, but is not continuously passable. Inside lava tubes, one can see lava stalactites, lava curbs, stacked tubes, bifurcating and coalescing channels, skylights, tube linings, and other features. Remelt features include submetallic appearance caused by a lack of gas bubbles in remelted material, soda straw like formations on the ends of lava stalactites, flowstone appearing linings, and small slumps.



Figure 8. Block Lava



Figure 9. Pahoehoe

Based on textural differences in lava flows within the Monument, many separate lava flows have been recognized. Some flows can begin as one flow type and change to another over time. Therefore, textural differences cannot be used as the sole method of separating flows. Aerial photographs and analysis of Landsat

images can also be used to distinguish flows from one another. Subtle differences in lava surfaces can often be detected on the images. These surface variances result from differing vegetation, weathering, and sediment coverage on individual flows (Lefebvre, 1977). But individual flows can also have uneven vegetation, differential weathering, and varying sediment coverage.

Most of the lava flows in the Monument are pahoehoe and were fed through tubes and tube systems, though there are some sheet flows. Structures representing both inflation and deflation of the lava surface can be seen along with both hot and cold collapses of lava tube roofs. Some lava flows produce tumuli (small mounds) or pressure ridges (elongate ridges) on their crusts. In some places squeeze-ups formed when pressure was sufficient to force molten lava up through tension fractures in the top of pressure ridges or cracks in the solidified crust of lava ponds (Fig.11). There are also pressure plateaus (Fig.12) that were produced by the sill-like injection of new lava beneath the crust of an earlier sheet flow that had not completely solidified.

Although it is not a convenient field method, geochemistry is the most accurate way to distinguish lava flows. The chemistry of the COM lava field has been studied in more detail than any other lava field in the ESRP. Geochemical examinations of the ESRP rocks have shown that the majority of the basalts to be olivine tholeiites. However, COM basalts are enriched in iron, phosphorus, titanium, and the alkali elements. Leeman, et al, (1976) believe the



Figure 10. Indian Tunnel lava tube.

magma that fed the COM eruptions evolved from the SRP tholeiites (silica-oversaturated basalt). This evolution could be the result of fractionation of the source magma (separation through crystallization) or crustal assimilation (melting and incorporation of crustal rocks as the magma migrated toward the surface of the earth) (Leeman, et al, 1976; Leeman, 1982; Kuntz, et al, 1992). For evolved lava, crustal contamination produces lava with silica (SiO_2) ranges of ~49% to 64%, while crystal fractionation produces lava with silica ranges of ~44% to 54% (Kuntz, et al, 1986).

Other lava features include spatter cones (Fig. 13) that formed when fluid globs (spatter) were ejected short distances (generally <200 ft or <60 m) from the vents and accumulated immediately around the

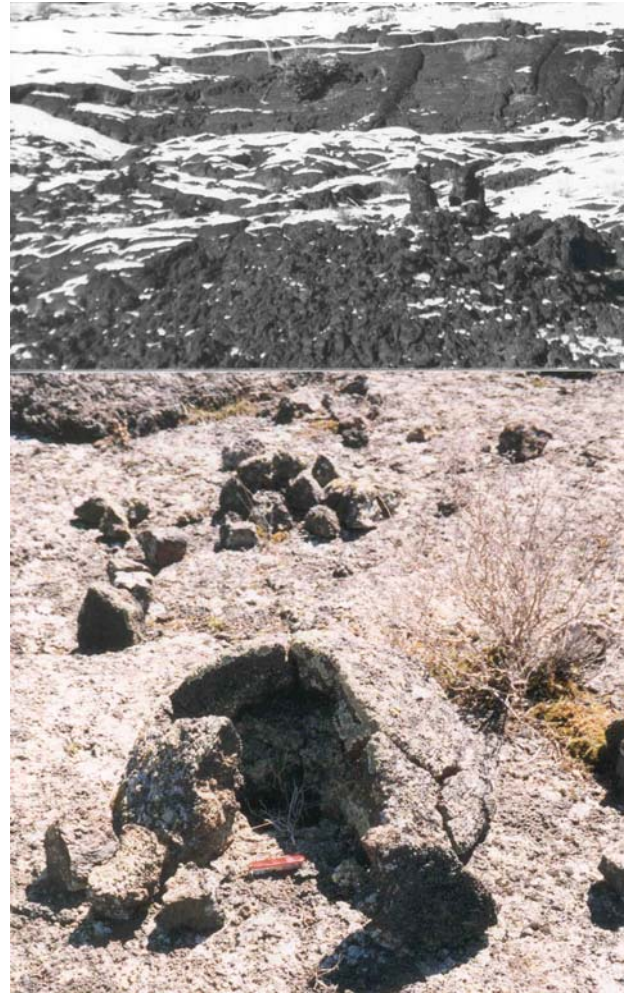


Figure 11. Squeeze-ups: Top—squeeze ups that came from the tension fracture in the top of a pressure ridge; Bottom—squeeze-up that looks like a candy kiss that oozed up through a crack in the crust of a lava pond.



Figure 12. Pressure Plateau on south side of Broken Top (lower right). (Also note slumping and open fissures on side of cone)

vent forming short steep-sided cones. Along eruptive fissures where a whole segment erupted, spatter can accumulate to produce low ridges called spatter ramparts. Hornitos, also known as rootless vents, are similar in appearance to spatter cones, but formed from spatter ejected from holes in the crust of a lava tube instead of directly from a feeding fissure. The Monument has collapse features known as sinks or pit craters (Fig. 14). During some eruptions, pieces of crater walls were carried off like icebergs by the lava flows. These wall chunks are known as rafted blocks (Fig. 15). Devils Orchard in northern part of the Monument is an entire field of rafted blocks that were carried off from the North Crater area. The Monument contains more than 580 Kipukas. Kipukas are older high areas that younger lava flowed around, but not over. They often appear as grassy hills surrounded by relatively barren lava.



Figure 13. Spatter Cone



Figure 14. Pit Craters
Caves

Besides shelly pahoehoe areas that contain many small open tubes and blisters and the numerous lava tubes associated with tube fed pahoehoe flows there are other kinds of caves found in the Monument. They include fissure caves associated with the Great Rift, many, such as Bear's Den waterhole, are ice floored. Flowing lava also can produce shallow caves and overhangs at flow fronts and as a result of the inflation process. Differential weathering of agglutinated cinders on some cinder cones has also generated a few shallow caves (Fig. 16); less firmly welded or sintered layers being more easily eroded. Some of these small caves are over 10 feet deep.



Figure 15. Rafted Blocks at Devil's Orchard



Figure 16. Differential weathering cave.

Stratigraphy and Dating

A stratigraphic section for the Monument consists of very few rock types. Surficial lithology is limited to the basaltic lava flows in all locations except the north end of the

monument. It is assumed that the rocks that underlie other parts of the ESRP, such as rhyolite flows, also underlie much of the Monument, though no drilling has been conducted. Unconsolidated sediments primarily include windblown silt (loess) and sand deposits, cinder deposits, alluvium along streams, colluvium at the base of steep slopes, and lacustrine deposits associated with ponds.

Dating methods used in the Monument include mapping field relations, magnetic polarity studies, dendrochronology, radiocarbon analyses, and K-Ar dating. In the past few decades, Kuntz, et al, (1988, 1989, 1989a, 1989b) and Champion, et al, (1989) have achieved detailed mapping of flows and their relationships with one another. The sequence of eruptions deduced from the superposition of volcanic landforms forms the base for other stratigraphic investigations.

The first efforts to determine the age of the lava flows in the Monument utilized dendrochronology (tree ring dating). One of the first attempts involved a limber pine tree known as the "Triple Twist Tree", which was growing in a crack on the North Crater Flow. It can still be seen today, but died back in 1961 and at the time of study already had rotted heartwood. The tree, about 16 inches (40 cm) in diameter, had 1,350 countable rings and was estimated to be 1,500 years old, allowing for the missing heartwood. The lava would be a minimum of 1,650 years old if it is assumed that it took at least 150 years for the soil to accumulate for it to grow in (Stearns, 1963). Dendrochronology is not applicable to the older flows because the trees despite their remarkable longevity are not long-lived enough, but at least tree ring dating helps establish a minimum age for the younger flows.

For a more precise age, absolute dating techniques that employ radioactive decay rates of various elements can be used. The decay of potassium to argon in the basalts of the Monument was analyzed by

Armstrong, et al, (1975). The results indicated that the basalts had a K-Ar age of zero million years, or are younger than the minimum datable age for this dating technique, which is approximately 10,000 to 20,000 years for rocks rich in potassium and probably 100-200 ka for SRP basalts. Recently researchers have been collecting samples for Argon 40/39 dating of the older flows.

Carbon-14 dating has been the most successful technique used to date Late Pleistocene and Holocene age flows in the Monument. Because lava flows obliterate nearly all carbon life during the eruptions, sources of carbon for dating are scarce. To find datable carbon for the first radiocarbon investigation, scientists dug beneath the lava flows at the perimeter of the lava field to uncover buried carbonized roots of the plants burned by the flowing lava (Bullard, 1970). What is believed to be sagebrush rootlets were found beneath a pahoehoe flow at the southern edge of the Blue Dragon flow. Analyses of two separate samples resulted in carbon-14 dates of $2,110 \pm 90$ years BP and $2,050 \pm 80$ years BP. Based on stratigraphy, it was estimated that this was one of the youngest flows in the COM lava field. Later investigations used both carbon retrieved from digging under the lava and small carbonized pieces of trees collected from tree molds in lava flows. Dating carbon from many localities throughout the Monument lava fields has yielded age dates ranging from approximately 15,000 to 1,700 years BP (Bullard, 1970; Kuntz, et al, 1982). Based on all dating techniques and stratigraphic investigations, it has been determined that COM lava field was formed during at least 8 eruptive periods/episodes between approximately 15,000 years to 2,000 years BP (Kuntz, et al, 1982, 1986, 1988, 1992). Within limits of analytical error, Kings Bowl and Wapi lava fields are contemporaneous and formed about 2,200 years ago.

Other Rock Units

There are no bore holes with cores to provide direct information about the rock

units underlying the basalt flows in the Monument. It is assumed that rhyolite exists at depth beneath the surface of the Monument, as it does elsewhere in the ESRP. Xenoliths of pumice, Challis volcanics, and granulite have been found in some of the basalts of the COM lava field. The pumice may be related to the rhyolitic eruptions that occurred within the ESRP, Challis volcanics, Mississippian sedimentary rocks, or Tertiary intrusives. The granulite is a metamorphic rock generally found in the cratonic basement and may be the rock type underlying the rhyolite.

The north end of the Monument is unlike the rest of the Monument and contains six sedimentary, volcanic, and intrusive rock units. Two intrusive-rock map units are exposed in outcrop in the north end of the Monument. Hornblende quartz monzonite is highly weathered and altered, weakly foliated, medium grained and equigranular. Plagioclase is the most abundant mineral in the monzonite and is accompanied by orthoclase, quartz and chloritized hornblende. Biotite granite is also exposed along the base of the Pioneer Mountains within the Monument. The granite contains quartz, biotite (altered to chlorite) and orthoclase (altered to the sericite). It also is highly weathered and is medium-coarse grained and equigranular in texture.

Eocene age Challis volcanics are present in the north end of the Monument and consist of welded tuff, lava flows interbedded with tuff breccia, and tuff breccia. The Challis tuff is an ash-flow deposit that overlies the tuff-breccia unconformably. The tuff ranges in color from light brownish gray to moderate orange pink with silica veins and lenses. The tuff breccia consists of lithic fragments (some of which are pumice), crystals and devitrified glass. These fragments were apparently derived from previously deposited or interbedded rhyodacite lava flows. Both units probably originated from ash flows and breccia flows issuing out of eruptive centers north of the monument boundary during Eocene time.

There are also Mississippian age sandstone, siltstone, claystone, and minor conglomerate of the Copper Basin Formation in the north end of the Monument. The conglomerate is gray with clasts to cobble size. The sandstone is very fine to fine grained, olive gray to medium gray, and sole marks are common in places. The siltstone is locally laminated and medium to dark gray. The claystone is dark gray to black, locally laminated, and contains pebbles of chert and quartzite in places.

Surficial deposits are the youngest materials mapped inside the Monument. The thickness of the sediments in the north end of the Monument ranges up to about 100 feet (30m) along some of the stream drainages based on the well logs from the Monument water wells. Eolian silts and sands mantle some of the older lavas and continue to be eroded, transported, and deposited, particularly after such events as fire. Recent fires in the Kings Bowl area freeing sediments of their anchors clearly demonstrated eolian processes in action, i.e., deflation, active ripple migration and formation and migration of small sand dunes. Cinders also can often be observed saltating on cinder cones on windy days, thus the cones are a landform in flux and ever changing.

Mineralization

The mineralization within Crystal Pit spatter cone is unique. This open chamber contains large quantities of the secondary sulfate minerals, gypsum, mirabilite, and jarosite, all of which seem to be scarce or absent in caves in other volcanic regions. Crystal Pit is a teardrop shaped cavity approximately 120 feet (36m) deep that most likely fed the spatter cone at the surface. Microcrystalline mirabilite ($\text{Na}_2\text{SO}_4 \cdot 10\text{H}_2\text{O}$) and crystalline gypsum ($\text{CaSO}_4 \cdot 2\text{H}_2\text{O}$) are abundant at the bottom of the pit, covering the walls and jarosite ($(\text{KFe}_3(\text{SO}_4)_2(\text{OH})_6)$) occurs as a loose powdery yellow material on the cave floor (Peck, 1974). The presence of such large quantities of these minerals in a volcanic

cave is at best unusual and their origin is still unknown. Peck (1974) proposed that the mirabilite and gypsum are likely capillary groundwater deposits. Another suggestion is that the minerals are precipitates from an underground lake in Crystal Pit. However, the rocks forming the bottom of the pit are porous and would not likely allow a lake to form. Thus, the origin of these minerals is still in debate.

A unique blue lava surface is found on a few flows in COM lava field. One flow has been named the Blue Dragon lava flow because of the color (Fig. 17). This pahoehoe lava flow extends southwest from the base of Big Craters and was first noted by Russell (1902). The broken lava surface seems to have a series of color layers to the naked eye. The outermost layer is a thin ($<5\mu\text{m}$) film that appears pale to deep blue. Beneath this layer is a 0.1 inch (3mm) thick blue colored glass layer that grades with depth into a brown glass. Thin sections of the flow surface appear brown in transmitted light but the outer blue glass appears blue in reflected light. Faye and Miller (1973) suggested that the blue color is as a result of an electron transfer between iron ions or iron and titanium ions due to oxidation that alters the light absorption in the rock surface. The reason for this blue color is not yet clearly understood, and awaits further research.

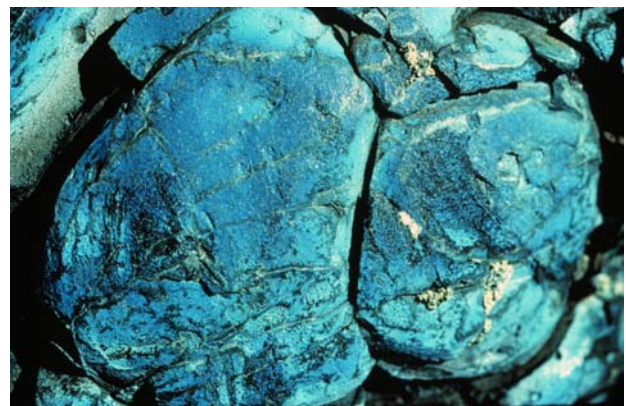


Figure 17. Blue Dragon

Paleontology

The igneous nature of the majority of the rocks in the Monument precludes a typical

fossil record. The exception is the sedimentary rocks of the Mississippian age Copper Basin Formation found on the northern edge of the Monument. The Copper Basin Formation is made up of interbedded claystone, siltstone, sandstone, and minor conglomerate. The trace fossil *Helminthoida* (believed to be burrows of a marine worm) is found on some bedding surfaces of the Copper Basin Formation. Because of the relative youth of the volcanism that dominates the Monument, fossils found within the lava fields must have accumulated since the volcanic activity, and therefore, are Pleistocene to Recent in age. They are primarily unaltered remains and trace fossils like tree molds.

Volcanic activity tends to destroy organic remains. However, tree molds are found in the lava flows of the ESRP. Tree molds are impressions in the solidified lava that form as trees are enveloped by the lava flows, begin to burn, release water and other vapors that quickly cool the surrounding lava, and leave behind a mold of the charred tree and occasionally some carbon residue. Generally, tree molds preserve impressions of the cracked, partly burnt wood but do not preserve bark or other textures that would aid in the identification of tree species. Tree molds can be both vertical (where the tree remained standing as it burned resulting in a columnar shaped hole in the lava) and horizontal (where the tree fell as it burned resulting in a linear mold in the lava). See figures 18 and 19. The deepest vertical tree mold mapped in the Monument to date is 2.9 meters (82 inches) and largest trunk width is 0.9 meters (35 inches). Some tree molds provide evidence of more than a dozen tree limbs. In the northern end of the Monument, more than 100 tree molds have been mapped. The two flows with the largest number of mapped tree molds are the Blue Dragon and Trench Mortar Flat flows.



Figure 18. Vertical Tree Mold

When lacking abundant sedimentary deposits to preserve the flora and fauna, the organic remains must be protected in some other way in order to survive over time. Lava tubes are commonly used by animals as hibernation, roosting, and den sites. They often provide a source of water and an escape from high temperatures. Animal bones accumulate in the tubes as inhabitants die naturally or are hunted and killed in the caves. Bones are also introduced into the caves as a result of human or animal disposal. Once in the cave, wind blown sediments may bury the bones, helping in the preservation process.

Exploration of such deposits in the lava tubes of the Snake River Plain has revealed bones of extinct animals, such as mammoth and camel, as well as modern large animals such as bear, wolf, bison, elk and pronghorn (Miller, 1989). Small animals identified mainly from regurgitated owl pellets include birds, reptiles, amphibians, snails and fish.



Figure 19. Horizontal Tree Mold

Although these animals may not have occupied the caves in life, they do offer some information about surrounding paleoecology. It should be noted that paleontological exploration of lava tubes on the ESRP has not been systematic and few have been within the Monument boundaries. Those caves, which have been excavated, were commonly archeological sites.

In addition to lava tubes, lava blisters have also accumulated a faunal record. The openings to lava blisters are generally small and drop to a floor, which can be 8-10 feet below the surface or more. This creates an excellent trap for larger animals that fall in and cannot escape. Generally, these animals are carnivores that most likely were lured into the trap by smaller prey such as a rabbit or a squirrel. Carnivores found in these blister traps on the ESRP include the now extinct noble marten, as well as other

animals no longer found in the area such as bison, wolverine, and Canada lynx (Miller, 1989). Although these traps contain a random collection of carnivores, they do not represent an accurate percentage of herbivores in relation to the carnivores, because herbivores are less likely to be lured into the trap.

A third type of unaltered fossil accumulation occurs in packrat nests. These nests, or middens, often contain twigs, leaves, pollen, cactus spines, porcupine quills and bones cemented by highly concentrated urine or "amber rat", which hardens and preserves the contents (Miller, 1989). These middens are an important contributor to the fossil record because of the ability to date the pollen and bone assemblages and relate that information to the paleoecology of the area.

Geologic Processes

In late summer of 2000, a Geoindicators Scoping Meeting was held at COM to determine what geologic processes are active within the Monument. Table 1 lists the geoindicators (a proxy for geologic processes) that are applicable to the Monument and indicates the relative ecological importance, human influence, and management significance of each geoindicator as rated by the scoping meeting work group. These geoindicators were adapted from Berger (1995). For the entire report see National Park Service, 2001.

Potential for Future Eruptions

We are at the end of the normal repose interval, the time of quiescence between eruptive periods. The COM lava field formed during eight eruptive periods with a recurrence interval averaging 2,000 years and it has been over 2,000 years since the last eruption. The constancy of the most recent output rates suggests that slightly over one cubic mile of lava will be erupted during the next eruption period. In the past, eruptions in the COM lava field have generally shifted to the segment of the Great Rift with the longest repose interval. Therefore, the next eruptive period should

Table 1. Geoindicators (proxies for geologic processes).

Geoindicator	Ecological Importance	Human Influence	Management Significance
Alpine and Polar			
Geological controls on perched water systems	H	L	M
Frozen ground activity (frost wedging)	L	L	H
Arid and Semi-Arid			
Desert microbiotic crusts and pavements	H	M*	M
Eolian processes	H	M	M
Groundwater			
Groundwater chemistry in the unsaturated zone	L	L	M
Groundwater level	H	M	M
Groundwater quality	L	L	H
Surface Water			
Surface water quality	M	M	H
Stream channel morphology	M	H	M
Streamflow	H	H	H
Wetlands extent, structure and hydrology	H	M	H
Hazards			
Volcanic unrest	H	L	H
Seismicity	L	L	L
Other (multiple environment)			
Soil and sediment erosion (water)	L	H	M
Soil compaction	L	M	M
Cave temperature and humidity regime	H	L	H
Hillslope processes	M	M	L
H – HIGHLY influenced by, or with important utility M – MODERATELY influenced by, or has some utility L – LOW or no substantial influence on, or utility * adjusted to reflect BLM input about grazing after the Monument expansion			

begin along the central portion of the Great Rift in the COM lava field, but may well propagate to the northern part of the monument in proximity to the loop road (Kuntz, et al, 1986). Initial flows, based on past history, will probably be relatively non-explosive and produce large-volume pahoehoe flows. Eruptions from potential vents on the northern part of the Great Rift

may be comparatively explosive and may produce significant amounts of tephra, destroy cinder cones by both explosion and collapse, and build new ones (Kuntz, et al, 1986). As yet, no comprehensive volcanic hazard assessment or plan has been done for the Monument.

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